

# SMART DA WITH JULES IN EMPIRE

Sanita Vetra-Carvalho,  
Tristan Quaife

With thanks to: Phil Browne



**National Centre for  
Earth Observation**  
NATURAL ENVIRONMENT RESEARCH COUNCIL



**University of  
Reading**

# Why data assimilation?

The aim of the data assimilation is to combine the prior knowledge or model forecast with likelihood or available observations primarily to:

- obtain analysis which is more accurate than either forecast or observations alone;
- learn about the model, e.g. model error, parameter values;
- assess observation impact.

# Why data assimilation?

Thus, given model forecast at time  $m$ :

$$\mathbf{x}^{(m)} = \mathcal{M}_m \left( \mathbf{x}^{(m-1)} \right) + \boldsymbol{\beta}^{(m)} \quad (1)$$

and observations at time  $m$ :

$$\mathbf{y}^{(m)} = \mathcal{H}_m \left( \mathbf{x}^{(m)} \right) + \boldsymbol{\beta}_{Ro}^{(m)}, \quad (2)$$

we aim to combine these two pieces of information.

There are three major ways of doing it:

- minimise a linear cost function - **variational DA**;
- linearly combine using Kalman gain & ens.- **ensemble DA**;
- solve Bayes theorem - **particle filters**.

# Why data assimilation?

Thus, given model forecast at time  $m$ :

$$\mathbf{x}^{(m)} = \mathcal{M}_m \left( \mathbf{x}^{(m-1)} \right) + \boldsymbol{\beta}^{(m)} \quad (1)$$

and observations at time  $m$ :

$$\mathbf{y}^{(m)} = \mathcal{H}_m \left( \mathbf{x}^{(m)} \right) + \boldsymbol{\beta}_{Ro}^{(m)}, \quad (2)$$

we aim to combine these two pieces of information.

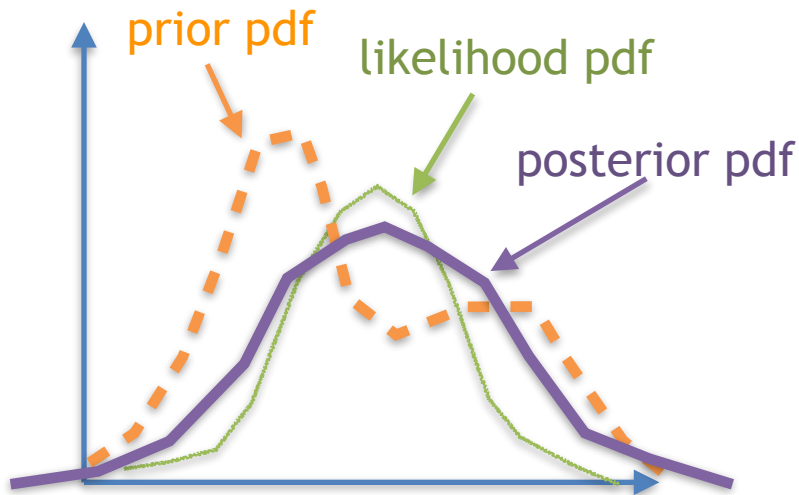
There are three major ways of doing it:

- minimise a linear cost function - **variational DA**;
- linearly combine using Kalman gain & ens.- **ensemble DA**;
- solve Bayes theorem - **particle filters**.

# Stochastic DA methods

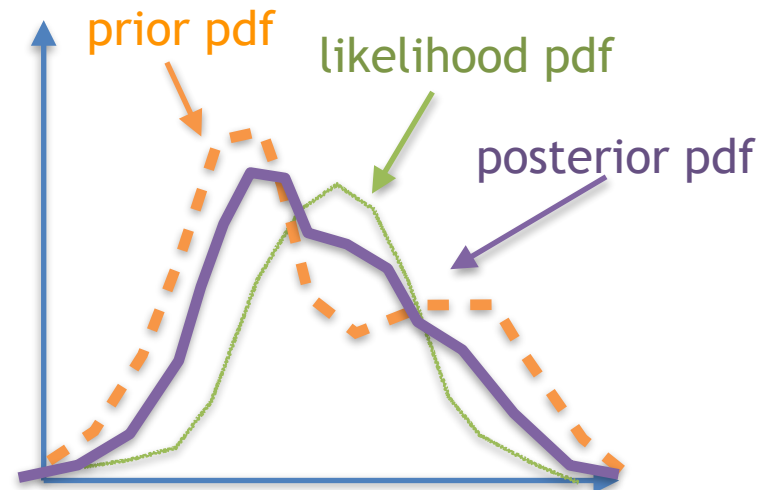
## Ensemble data assimilation (EnDA)

- based on Kalman Filter equations
- uses Gaussian assumptions of errors
- needs knowledge of prior errors,  $P$
- localisation needed for high-dimensional systems



## Particle filters (PF)

- based on Bayes theorem
- fully non-linear methods
- needs knowledge of model errors,  $Q$  but not prior errors
- curse of high-dimensionality



# Data assimilation in JULES

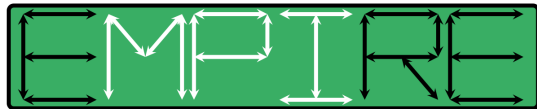
- **ADJULES (U Exeter):** provides adjoint based optimisation for finding JULES parameters
- **CARDAMOM (U Edinburgh):** retrieves terrestrial carbon cycle variables using obs. and model
- **PECAN (U Reading w/ Boston):** predictive ecosystem analyser with JULES added in 2016/17
- **JULES/EMPIRE (U Reading):** uses MPI to communicate between JULES and EMPIRE

# Data assimilation in JULES

- **ADJULES (U Exeter)**: provides adjoint based optimisation for finding JULES parameters
- **CARDAMOM (U Edinburgh)**: retrieves terrestrial carbon cycle variables using obs. and model
- **PECAN (U Reading w/ Boston)**: predictive ecosystem analyser with JULES added in 2016/17
- **JULES/EMPIRE (U Reading)**: uses MPI to communicate between JULES and EMPIRE

# EMPIRE

**EMPIRE** - Employing Message Passing Interface for Researching Ensembles



[www.met.reading.ac.uk/~darc/empire](http://www.met.reading.ac.uk/~darc/empire)

EMPIRE is: a system which easily connects any dynamical model to a DA method.

EMPIRE uses: MPI calls to setup communication and transfer data between the model and DA methods.

EMPIRE is: continuously developed at UoR



# EMPIRE

## Models connected to EMPIRE

<b>Model Name</b>	<b>Description</b>
HadCM3	Coupled ocean-atmosphere climate model
UM vn8.2	UK Met Office atmospheric forecast model configured as N512L70 global model
TELEMAC-2D	Unstructured finite element code, used for storm surge modelling
DALECv2	Land surface model
JULES	Operational land surface model
GOTM-ERSEM	Coupled hydrodynamic-biogeochemical model
NEMO	Ocean model
Liley-Bojak model	Neurofield model of electrocortical activity in the brain
Barotropic vorticity model	Simple 2D fluid dynamics model of barotropic flow on a torus
ECMWF IFS Single column model	A single column, coupled ocean-atmosphere model
MITgcm	Numerical model designed for study of the atmosphere, ocean, and climate
Enlil	A time-dependent 3D magnetohydrodynamics model of the heliosphere

# EMPIRE

Currently available  
DA methods are:

- LETKF
- EWPF
- IEWPF
- SIR PF
- 4DEnVar
- 3DVar

Future:

- More flavours of EnKF
- Variational methods

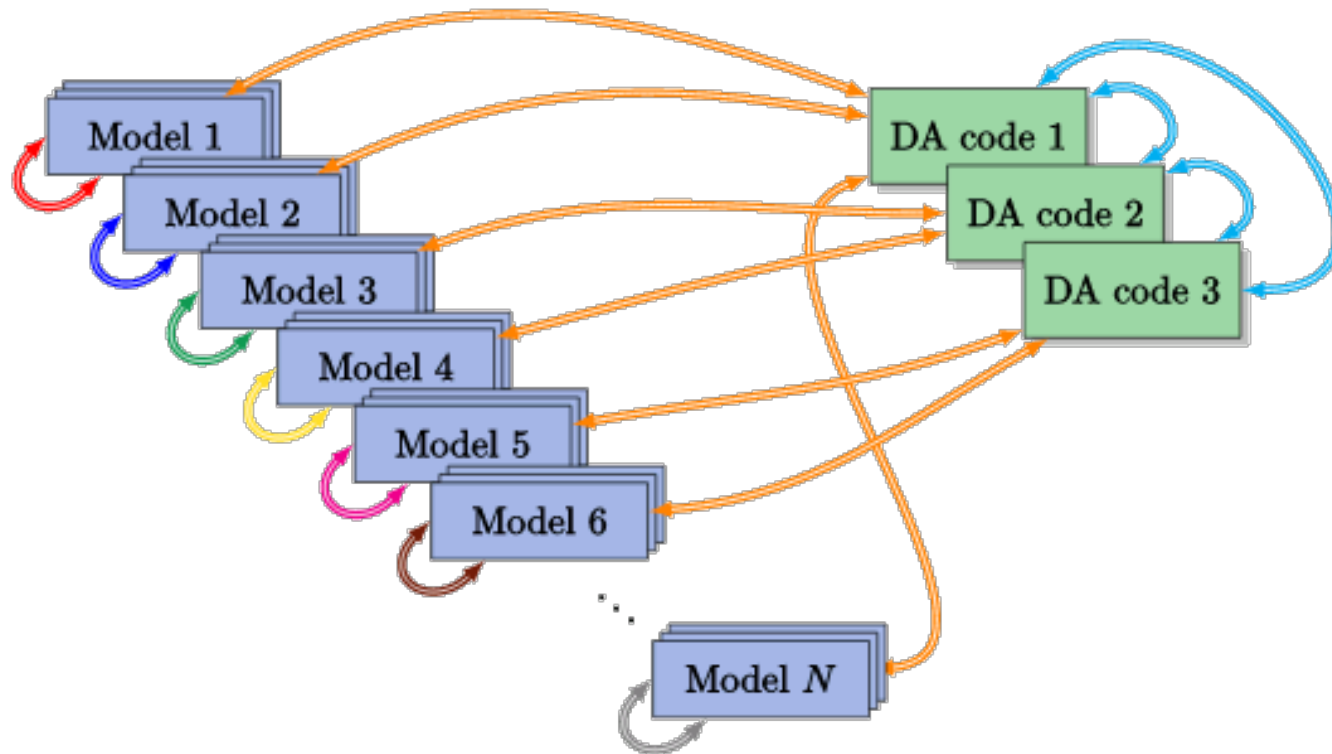


# EMPIRE

**Benefits of using EMPIRE** are obvious for both model developer and research communities:

- Easy to test your model with various DA methods;
- Easy to test your DA method with various models;
- No need to hack the model code to add DA;
- Quick way to decide which DA method will perform best for a given model;
- Little time spent on coupling a given model to EMPIRE; No need to change a given model flow and working;
- No need to spend time coding DA methods.

# EMPIRE



# EMPIRE

User needs to provide following operators to EMPIRE:

- $H, H^T$
- $R^{1/2}, R^{-1}, R^{-1/2}$
- $Q, Q^{1/2}$
- Distance between element of state vector and observation  
 $(HQH^T + R)^{-1}$

$EWPF, LETKF$ , twin experiment observation generation

# EMPIRE

To couple EMPIRE to JULES we added **ONE EMPIRE module** into the code - **empire\_module**, which includes subroutines to:

- Create MPI communication between JULES and EMPIRE
- Pack and unpack JULES state variables into 1D state vector for use in EMPIRE
- Send and receive data between JULES and EMPIRE
- Clean up MPI communication at the end of the run.

# EMPIRE

## Changes in JULES:

- the only modified routine in JULES is jules.F90
- preprocess using key EMPIRE (e.g. `#if defined(EMPIRE)`) - added in FCM code if used
- Routines called in jules.F90 are:
  - before time step loop: *mpi\_init(error)* & *empire\_initialise\_state()*
  - in the time step loop: *empire()*
  - at the end of the run: *empire\_cleanup()*

# EMPIRE

## Changes in JULES:

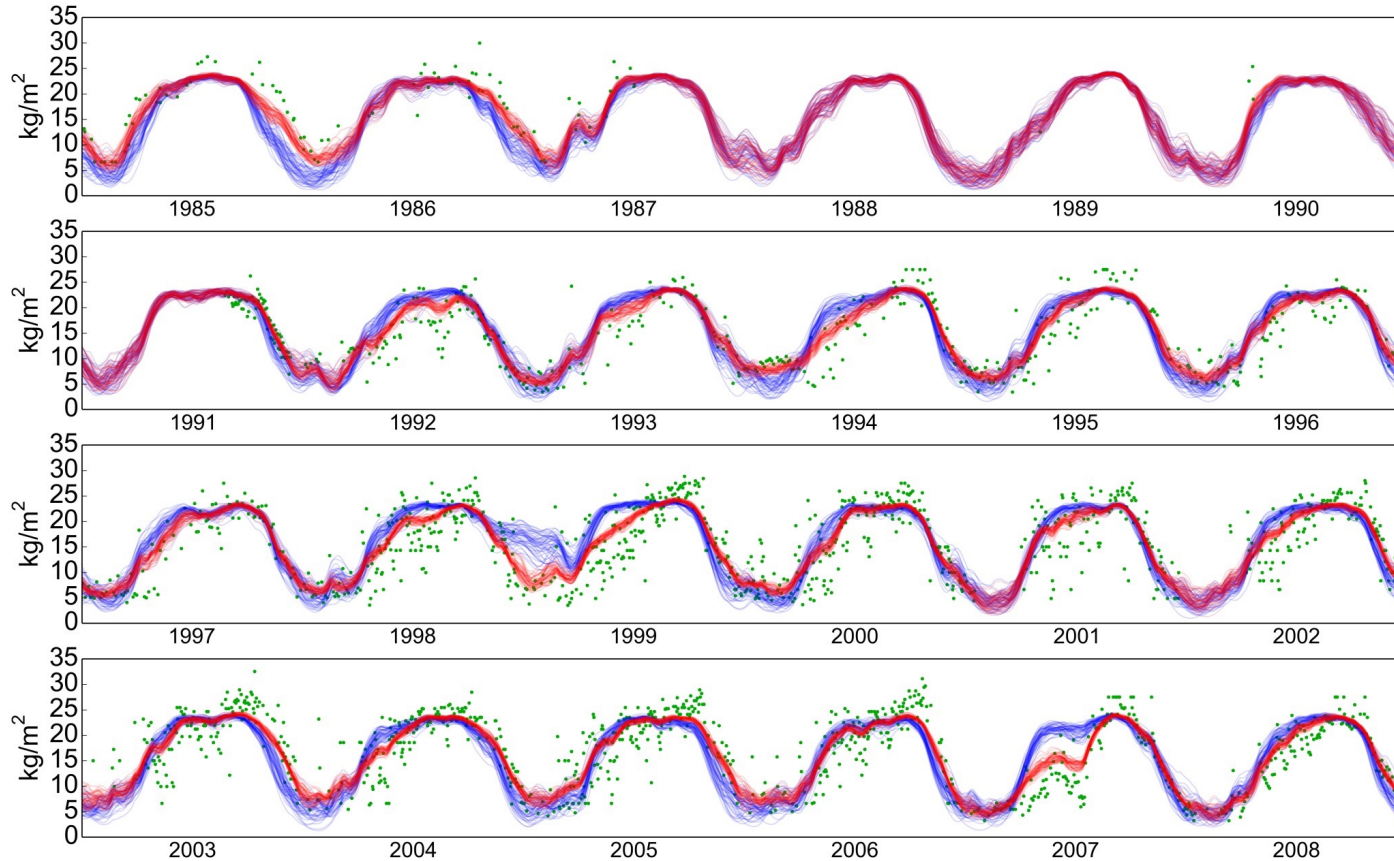
- the only modified routine in JULES is jules.F90
- preprocess using key EMPIRE (e.g. `#if defined(EMPIRE)`) - added in FCM code if used
- Routines called in jules.F90 are:
  - before time step loop: *mpi\_init(error)* & *empire\_initialise\_state()*
  - in the time step loop: *empire()*
  - at the end of the run: *empire\_cleanup()*

**Preparing patch to automate the process!!!**



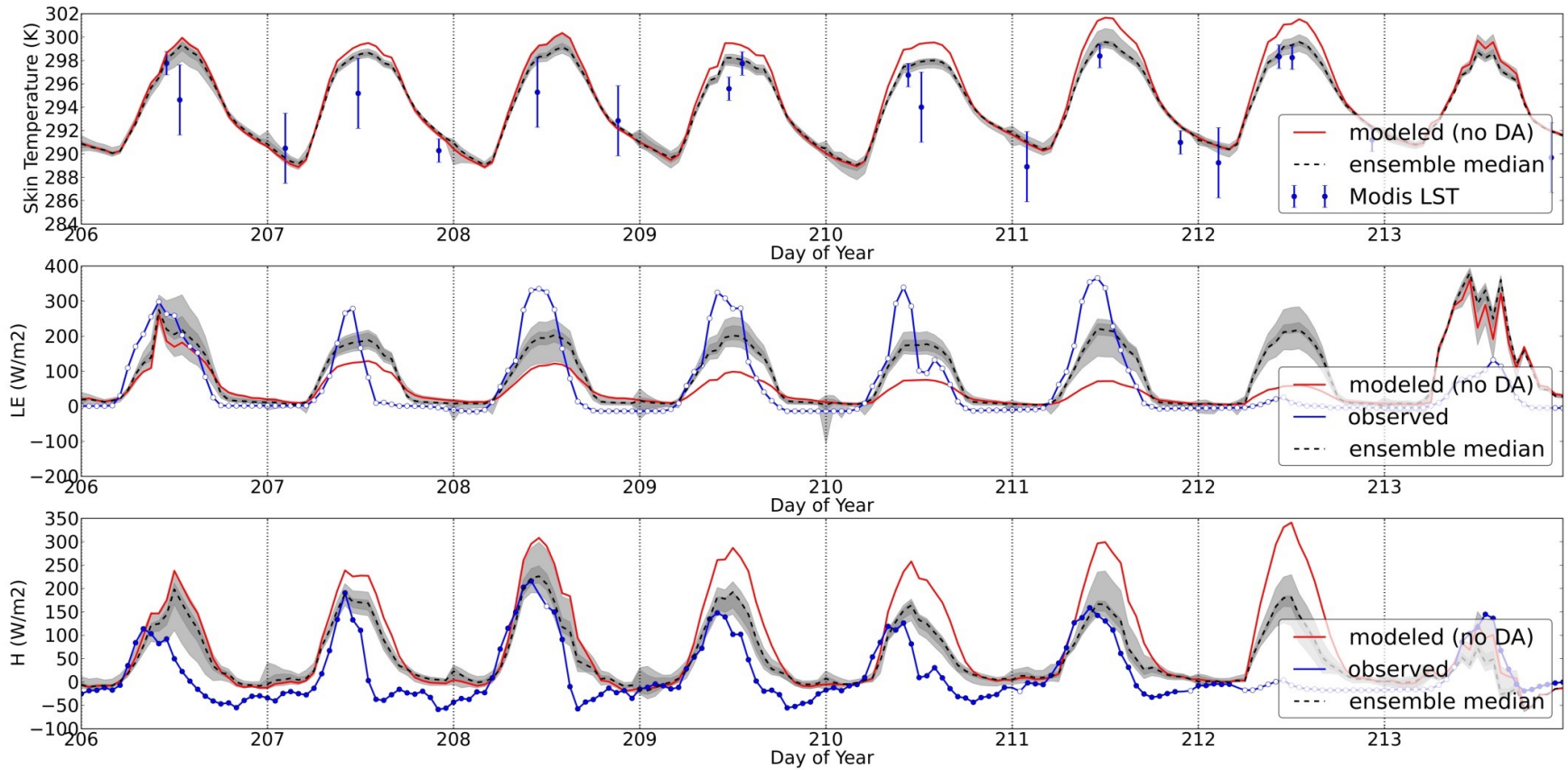
# Preliminary results

# Assimilation 30 years of ESA SM data (plus stochastic forcing from TAMSAT)



N = 50  
Site = Gana  
DA = ETKF  
obs = top soil moist.  
blue = no DA  
red = DA

# Assimilating MODIS LST & albedo



# Summary

- Work has just started
- JULES has been coupled with EMPIRE
- Positive impact of DA in preliminary results

## **Future:**

- Provide patches to users to couple their JULES code to EMPIRE (over next few months)
- Provide some basic operators for the user
- Research work in DA and JULES
- Ongoing work on EMPIRE with more DA methods added (not me)

# Questions?

SANGOMA website: [www.data-assimilation.net](http://www.data-assimilation.net)

EMPIRE website: [www.met.reading.ac.uk/~darc/empire](http://www.met.reading.ac.uk/~darc/empire)

Eager to learn how community  
is interested to use DA in JULES  
[s.vetra-carvalho@reading.ac.uk](mailto:s.vetra-carvalho@reading.ac.uk)